

BPS D-branes from an unstable D-brane in a curved background

Chanju Kim*

*Department of Physics, Ewha Womans University, Seoul 120-750, Korea*Yoonbai Kim,[†] Hwang-hyun Kwon,[‡] and O-Kab Kwon[§]*BK21 Physics Research Division and Institute of Basic Science, Sungkyunkwan University, Suwon 440-746, Korea*

(Received 8 June 2005; published 13 July 2005)

We find exact tachyon kink solutions of Dirac-Born-Infeld–type effective action describing an unstable D5-brane with a world volume gauge field turned on in a curved background. The background of interest is the ten-dimensional lift of the Salam-Sezgin vacuum and, in the asymptotic limit, it approaches $\mathbb{R}^{1,4} \times \mathbb{T}^2 \times \mathbb{S}^3$. The solutions are identified as composites of lower-dimensional D-branes and fundamental strings, and, in the BPS limit, they become a D4D2F1 composite wrapped on $\mathbb{R}^{1,2} \times \mathbb{T}^2$ where \mathbb{T}^2 is inside \mathbb{S}^3 . In one class of solutions we find an infinite degeneracy with respect to a constant magnetic field along the direction of the Neveu-Schwarz–Neveu-Schwarz field on \mathbb{S}^3 .

DOI: [10.1103/PhysRevD.72.026006](https://doi.org/10.1103/PhysRevD.72.026006)

PACS numbers: 11.25.Uv, 11.27.+d

I. INTRODUCTION

The study of unstable D-branes in string theory has led to a deeper understanding of the theory in various aspects [1]. In a dynamical aspect, it provides an example of time-dependent string background that can be described by the language of world sheet conformal field theory [2]. Descent relations among BPS and non-BPS D-branes of various dimensions provided a new perspective on the characteristics of D-branes and helped the development of the classification of D-brane charges in terms of K theory [3]. These rolling tachyons and tachyon solitons can also be dealt, at least qualitatively, in terms of Dirac-Born-Infeld (DBI)-type effective action [4,5].

Most of the studies so far have been performed on flat unstable D-branes. The purpose of the paper is an attempt to extend the analysis of unstable D-branes to a curved bulk background and find tachyonic kink solutions on them. The background that we consider is the ten-dimensional embedding of the supersymmetric vacuum, $\mathbb{R}^{1,3} \times \mathbb{S}^2$, of the Salam-Sezgin model [6]. The model is a gauged $\mathcal{N} = (1, 0)$ supergravity in six dimensions coupled to a tensor and an Abelian vector multiplet [7]. The vacuum solution is the unique nonsingular solution of the model with maximal four-dimensional spacetime symmetry. It can be embedded into ten dimensions as the type-IIA supergravity background solution, $\mathbb{R}^{1,3} \times \mathbb{S}^2 \times \mathcal{H}^{2,2} \times \mathbb{S}^1$, where $\mathcal{H}^{2,2}$ represents a three-dimensional hyperboloid. In view of recent works on the dynamics of D-branes in NS5-brane background initiated by Kutasov [8], it is interesting to notice that [6], in the asymptotic limit at large distances, the local geometry of

the background of Salam-Sezgin vacuum approaches the NS5-brane near-horizon geometry [9,10]. However, there is a difference in that the string coupling constant goes to zero in the asymptotic limit of the background, while it blows up in the throat region of the NS5-brane. In this limit, it is valid to study non-BPS D-branes in terms of DBI-type effective theory. Still our work may have an implication for recent developments along the line of [8].

In the asymptotic limit, the background approaches $\mathbb{R}^{1,4} \times \mathbb{T}^2 \times \mathbb{S}^3$. The object of consideration is a non-BPS D5-brane whose world volume lies on $\mathbb{R}^{1,2} \times \mathbb{S}^3$. After the unstable D-brane decays, one may expect the generation of lower-dimensional stable brane configurations [5,11] and emission of energy to closed string degrees of freedom [12,13]. The former, specifically the codimension-one object, is what we would like to study in this paper. We find two classes of exact solutions both of which are identified as thick D4-branes on $\mathbb{R}^{1,2} \times \mathbb{T}^2$ where the two-torus is embedded in the three-sphere. One class of the solutions contains fundamental string and two D2-branes; one is tubular and the other flat. The flat D2-brane disappears when a constant magnetic field, h , along the direction of the Neveu-Schwarz–Neveu-Schwarz (NS-NS) field on \mathbb{S}^3 vanishes. In the thin limit, the solution becomes BPS when $h = 0$ and the energy expression is given by a BPS sum rule. The other class involves one tubular D2-brane and fundamental string. In the thin limit it becomes essentially identical to the former class of the solutions except that it has infinite degeneracy with respect to h .

The rest of this paper is organized as follows. In Sec. II, we give a brief review of the bulk background on which we consider tachyon kink solutions of an unstable D5-brane. In Sec. III, by considering an unstable D5-brane in the given background, we obtain stable codimension-one D-brane configurations and identify their BPS limit. We conclude the paper in Sec. IV.

*Electronic address: cjkim@ewha.ac.kr†Electronic address: yoombai@skku.edu‡Electronic address: hkkwon@skku.edu§Electronic address: okab@skku.edu

II. BACKGROUND GEOMETRY

In this section we briefly describe the background of the ten-dimensional lift of Salam-Sezgin vacuum¹ and set the convention. The type-IIA supergravity action is given by

$$S_{\text{NS}} = \frac{1}{2\kappa_{10}^2} \int d^{10}x \sqrt{-G} e^{-2\Phi} \left(R + 4\partial_\mu \Phi \partial^\mu \Phi - \frac{1}{2 \cdot 3!} H^{\mu\nu\rho} H_{\mu\nu\rho} \right) \quad (1)$$

for the massless NS-NS fields in the string frame. Here κ_{10} is the ten-dimensional gravitational coupling constant, Φ is

$$ds^2 = e^{-\phi/2} ds_6^2 + \frac{1}{2g^2} \left[d\rho^2 + \frac{\cosh^2 \rho}{\cosh 2\rho} (d\sigma - gA)^2 + \frac{\sinh^2 \rho}{\cosh 2\rho} (d\tau + gA)^2 \right] + d\chi^2, \\ H = \frac{\sinh \rho \cosh \rho}{g^2 (\cosh 2\rho)^2} d\rho \wedge (d\sigma - gA) \wedge (d\tau + gA) + \frac{1}{2g \cosh 2\rho} dA \wedge [\cosh^2 \rho (d\sigma - gA) - \sinh^2 \rho (d\tau + gA)] + dB', \quad (2) \\ e^\Phi = (\cosh 2\rho)^{-1/2} e^{-\phi/2},$$

where g is a rescaled coupling constant for the gauge potential A , (ρ, σ, τ) parametrize the hyperboloid $\mathcal{H}^{2,2}$, and χ the circle S^1 .

Using the ansatz (2), any solution of the Salam-Sezgin model can be lifted to ten dimensions. The one of our interest is the supersymmetric vacuum solution, $\mathbb{R}^{1,3} \times S^2$, with the magnetic monopole flux on S^2 , which is given by

$$ds_6^2 = dx_4^2 + \frac{1}{8g^2} (d\theta^2 + \sin^2 \theta d\varphi^2), \\ A = -\frac{1}{2g} \cos \theta d\varphi, \quad dB' = 0, \quad \phi = 0, \quad (3)$$

where dx_4^2 represents the line element of the four-dimensional Minkowski space. Inserting the solution (3) into the ansatz (2) gives the background of the ten-dimensional lift of the Salam-Sezgin vacuum. In the large- ρ limit, the embedded solution simplifies and can be written as

$$ds^2 = dx_6^2 + \frac{1}{2g^2} d\rho^2 + \frac{1}{8g^2} [d\theta^2 + \sin^2 \theta d\varphi^2 + (d(\sigma - \tau) + \cos \theta d\varphi)^2], \\ H = \frac{1}{8g^2} \sin \theta d\theta \wedge d\varphi \wedge d(\sigma - \tau), \quad \Phi = -\rho, \quad (4)$$

where $0 \leq \theta \leq \pi$, $0 \leq \varphi \leq 2\pi$, $0 \leq (\sigma - \tau) \leq 4\pi$, and dx_6^2 represent the line element of $\mathbb{R}^{1,3} \times T^2$. Besides the fact that there is a linear dilaton background along the

¹The details of the construction of the background can be found in [6].

the dilaton field, and $H_{\mu\nu\rho}$ is the field strength of NS-NS two-form potential $B_{\mu\nu}$. Ramond-Ramond (R-R) and fermionic fields do not play a role for the embedding and may be set to zero.

The six-dimensional Salam-Sezgin model is a gauged $\mathcal{N} = (1, 0)$ supergravity coupled to a tensor and an Abelian vector multiplet. Its bosonic sector consists of the metric, real scalar field, ϕ , Abelian one-form and two-form gauge fields, A and B' . Through a chain of dimensional reductions and truncations, one can see that the model can be obtained by dimensionally reducing IIA supergravity on $\mathcal{H}^{2,2} \times S^1$. Here $\mathcal{H}^{2,2}$ is the quadric, $\mu_1^2 + \mu_2^2 - \mu_3^2 - \mu_4^2 = 1$, in Euclidean space \mathbb{R}^4 . The embedding is described by the ansatz

direction \mathbb{R}_ρ , one can notice that the coordinates $(\theta, \varphi, \sigma - \tau)$ parametrize S^3 as the Hopf fibration of S^1 over S^2 and the field strength H is proportional to the volume form of the unit three-sphere. Therefore, the background (4) is locally identical to the near-horizon geometry of the NS5-brane [9,10]. However, as remarked in the previous section, the behavior of the dilaton is opposite in the respective limits of the two background solutions.

In the following discussions, it turns out to be useful to rescale the angular variables by the radius of S^3 , $R = 1/\sqrt{8g^2}$, to the variables (u, v, w) and use a gauge-fixed value of the gauge potential B , which brings (4) to

$$ds^2 = dx_6^2 + 4R^2 d\rho^2 + du^2 + \sin^2 \left(\frac{u}{R} \right) dv^2 + \left[dw + \cos \left(\frac{u}{R} \right) dv \right]^2, \quad (5)$$

$$B = -\cos \left(\frac{u}{R} \right) dv \wedge dw, \quad (6)$$

$$\Phi = -\rho. \quad (7)$$

III. D4D2F1-COMPOSITES FROM UNSTABLE D5-BRANE

In this section we will study tachyon kink solutions on an unstable D5-brane in the large- ρ limit of the ten-dimensional lift of Salam-Sezgin vacuum on $\mathbb{R}^{1,3} \times T^2 \times \mathbb{R}_\rho \times S^3$ described by Eqs. (5)–(7). We consider the D5-

brane on $\mathbb{R}^2 \times \mathbb{S}^3$ with the coordinates (z, ρ, u, v, w) where z is one of the spatial coordinates² of $\mathbb{R}^{1,3}$. The string coupling constant goes to zero in the limit and many features of the dynamics of unstable D-branes can be described by the DBI-type world volume effective action [14]

$$S_{D5} = -\mathcal{T}_5 \int d^6 \xi e^{-\Phi} V(T) \times \sqrt{-\det(g_{\mu\nu} + F_{\mu\nu} + B_{\mu\nu} + \partial_\mu T \partial_\nu T)}, \quad (8)$$

where \mathcal{T}_5 is the tension of the non-BPS D5-brane, $F_{\mu\nu}$ is the field strength of the U(1) world volume gauge field, $T(x)$ is the real tachyon field, and ξ^μ ($\mu = 0, \dots, 5$) runs over the coordinates (t, z, ρ, u, v, w) . The background metric $g_{\mu\nu}$ and NS-NS field $B_{\mu\nu}$ are given by the pullbacks of (5) and (6) on the D5-brane. For the tachyon potential $V(T)$, any runaway potential with $V(0) = 1$ and $V(\pm\infty) = 0$ is allowed for the existence of the D-brane configuration of our interest that is consistent with universal behavior in tachyon condensation [5], but we assume a specific form for exact solutions [15]

$$V(T) = \frac{1}{\cosh\left(\frac{T}{R}\right)}, \quad (9)$$

where R turns out to be identical to the compactification scale R in (5) for the kink solutions we find.

For the rest of the paper we will study codimension-one solutions of (8) under the ansatz

$$T = T(u, w), \quad (10)$$

and the nonvanishing components of the world volume gauge field strength are

$$F_{0z} = E_z(u, w), \quad F_{vz} = \alpha(u, w), \quad F_{vw} = h(u, w). \quad (11)$$

It turns out to be an appropriate ansatz for the tachyon and the U(1) gauge field to support a tachyon tube embedded in a codimension-one D4-brane.³ With this ansatz, the coordinate ρ decouples from the others and hence the presence of the dilaton background (7) plays no role. Therefore, in the following, we will simply ignore the dilaton background in the action.

Applying the Bianchi identity

$$\partial_\mu F_{\nu\lambda} + \partial_\nu F_{\lambda\mu} + \partial_\lambda F_{\mu\nu} = 0, \quad (12)$$

²Inclusion of the linear dilaton coordinate ρ in the world volume of D5 is not essential for the type of solutions we consider; it can be trivially replaced by any other spatial coordinate on $\mathbb{R}^{1,3}$ without changing the rest of the analysis.

³One can also consider solutions with all w replaced by v in (10) and (11). Since the coordinates v and w are symmetric (except minus signs in vw components of two-form fields and the range of the variables), the resulting solutions will be essentially identical to the present case.

we can further simplify the ansatz (11). The result is

$$E_z = \text{const}, \quad \alpha = \text{const}, \quad h = h(w). \quad (13)$$

Substituting the nonvanishing gauge field (13), the tachyon (10), and the NS-NS two-form field (6) into the action (8), we have

$$S_{D5} = -\mathcal{T}_5 \int d^6 \xi V(T) \left\{ (1 - E_z^2 + \alpha^2) [1 + (\partial_u T)^2 + (\partial_w T)^2] + (1 - E_z^2) [1 + (\partial_u T)^2] \times \left[\left(h - \cos\left(\frac{u}{R}\right) \right)^2 - \cos^2\left(\frac{u}{R}\right) \right]^{1/2} \right\}. \quad (14)$$

As has been done for every tubular object, we take a critical electric field along the z direction, $|E_z| = 1$, as a basic ansatz, and then the action (14) takes a simple form,

$$S_{D5} = -\mathcal{T}_5 \alpha \int d^6 \xi V(T) \sqrt{1 + (\partial_u T)^2 + (\partial_w T)^2}. \quad (15)$$

Then the equations of motion for the tachyon and gauge field are given by

$$\partial_u \left[\frac{\alpha V}{\sqrt{-X}} \partial_u T \right] + \partial_w \left[\frac{\alpha V}{\sqrt{-X}} \partial_w T \right] = \frac{\sqrt{-X}}{\alpha} \frac{dV}{dT}, \quad (16)$$

$$\partial_u \left[\frac{\alpha V}{\sqrt{-X}} h \partial_u T \partial_w T \right] = \partial_w \left[\frac{\alpha V}{\sqrt{-X}} h (1 + (\partial_u T)^2) \right], \quad (17)$$

$$\partial_u \left[\frac{\alpha V}{\sqrt{-X}} \cos\left(\frac{u}{R}\right) \partial_u T \partial_w T \right] = \partial_w \left[\frac{\alpha V}{\sqrt{-X}} \cos\left(\frac{u}{R}\right) \times (1 + (\partial_u T)^2) \right], \quad (18)$$

where $X = -\alpha^2 [1 + (\partial_u T)^2 + (\partial_w T)^2]$. It is not difficult to solve these equations. After some manipulations with (16) and (18) we find that they are consistent only when

$$\partial_u T \partial_w T = 0, \quad (19)$$

i.e., the tachyon field T is a function of either u or w but not both. Then from (17) and (18) it is easy to see that h should be a constant.

A. $T = T(u)$

We first consider the case $\partial_w T = 0$, i.e., T depends only on u . Then the only nontrivial equation is (16) (with $\partial_w T = 0$). This equation is actually exactly the same as that for the usual tachyon tube [16,17] in lower dimensions and can also be directly derived from the reduced action

$$S_{D5} = -\mathcal{T}_5 \alpha \int d^6 \xi V(T) \sqrt{1 + \left(\frac{dT}{du}\right)^2}. \quad (20)$$

It is now quite straightforward to solve Eq. (16). Rewriting the equation, we find

$$\partial_u \left[\frac{V(T)}{\sqrt{-X}} \right] = 0. \quad (21)$$

Thus the whole equations of motion reduce to a single first-order differential equation

$$\beta = \frac{\mathcal{T}_5 V(T)}{\sqrt{-X}}, \quad (22)$$

where β is an integration constant. With the tachyon potential (9), we can find the exact solution in a closed form,

$$\sinh\left(\frac{T(u)}{R}\right) = \pm \sqrt{\left(\frac{\mathcal{T}_5}{\alpha\beta}\right)^2 - 1} \cos\left(\frac{u}{R}\right). \quad (23)$$

The energy-momentum tensor $T^{\mu\nu}$ of the system is given by

$$T^{\mu\nu} = \frac{\mathcal{T}_5 V}{\sqrt{-X}\sqrt{-g}} C_s^{\mu\nu}, \quad (24)$$

where $C_s^{\mu\nu}$ is the symmetric part of the cofactor of $X_{\mu\nu} = g_{\mu\nu} + F_{\mu\nu} + B_{\mu\nu} + \partial_\mu T \partial_\nu T$. For the solution (23), nonvanishing components are

$$\begin{aligned} T^{uu} &= \frac{1}{\sin(u/R)} \left[1 + \alpha^2 + h^2 - 2h \cos\left(\frac{u}{R}\right) \right] \Sigma(u), \\ T^{zz} &= -\frac{1}{\sin(u/R)} \left[1 + h^2 - 2h \cos\left(\frac{u}{R}\right) \right] \Sigma(u), \\ T^{uu} &= -\frac{\beta\alpha^2}{\sin(u/R)}, \quad T^{ww} = -\frac{\alpha^2}{\sin(u/R)} \Sigma(u), \quad (25) \\ T^{tv} &= -\frac{\alpha}{\sin(u/R)} \Sigma(u), \quad T^{tw} = \alpha \cot\left(\frac{u}{R}\right) \Sigma(u), \\ T^{zw} &= \frac{\alpha}{\sin(u/R)} \left[h - \cos\left(\frac{u}{R}\right) \right] \Sigma(u), \end{aligned}$$

where

$$\begin{aligned} \Sigma(u) &= \beta[1 + (\partial_u T)^2] = \frac{1}{\beta\alpha^2} (\mathcal{T}_5 V)^2 \\ &= \frac{\mathcal{T}_5^2 / \beta\alpha^2}{\left[\left(\frac{\mathcal{T}_5}{\beta\alpha}\right)^2 - 1\right] \cos^2\left(\frac{u}{R}\right) + 1}. \end{aligned} \quad (26)$$

The electric flux Π_i , which is the conjugate momentum of A_i , is calculated as

$$\Pi_i = \frac{\mathcal{T}_5 V}{\sqrt{-X}} C_A^{0i}, \quad (27)$$

where C_A^{0i} is the antisymmetric part of the cofactor of $X_{\mu\nu}$. Then the solution (23) has two nonzero components,

$$\begin{aligned} \Pi_z &= \left[1 + h^2 - 2h \cos\left(\frac{u}{R}\right) \right] \Sigma(u), \\ \Pi_w &= -\alpha \left[h - \cos\left(\frac{u}{R}\right) \right] \Sigma(u). \end{aligned} \quad (28)$$

Note that, except T^{uu} , all nonzero components of $T^{\mu\nu}$ and Π_i depend on $\Sigma(u)$ which has a peak at $u = \pi R/2$ where the NS-NS two-form background field (6) vanishes. In fact, in the limit $\beta \rightarrow 0$ it becomes a delta function

$$\Sigma(u) \rightarrow \frac{\pi R \mathcal{T}_5}{\alpha} \delta(u - \pi R/2), \quad (29)$$

and so do the nonvanishing components of $T^{\mu\nu}$ and Π_i (T^{uu} goes to zero in this limit). Therefore, when β is small, the coordinate u is essentially fixed at $\pi R/2$ and the solution represents a dimensionally reduced configuration. From the background metric (5), we see that the configuration spans T^2 in the three-sphere:

$$\begin{aligned} ds_{S^3} &= du^2 + \sin^2\left(\frac{u}{R}\right) dv^2 + \left[dw + \cos\left(\frac{u}{R}\right) dv \right]^2 \xrightarrow{\frac{u}{R} \rightarrow \frac{\pi}{2}} ds_{T^2} \\ &= dv^2 + dw^2. \end{aligned} \quad (30)$$

Of course, this is because the three-sphere with the above metric is a Hopf fibration of S^2 with coordinates (u, v) and hence the circle $u = \pi R/2$ along the equator of the two-sphere corresponds to T^2 in the three-sphere.

The coupling to the bulk R-R fields can be read off from the Wess-Zumino term for unstable D-branes [5,18],

$$S_{\text{WZ}} = \mathcal{T}_5 \int V(T) dT \wedge C_{\text{RR}} \wedge e^{F+B}. \quad (31)$$

For the solution (23),

$$dT = \partial_u T du, \quad (32)$$

$$F + B = dt \wedge dz + \alpha dz \wedge dv + \left[h - \cos\left(\frac{u}{R}\right) \right] dv \wedge dw.$$

Then, in the thin limit $\beta \rightarrow 0$ we can use (29) to simplify S_{WZ} ,

$$\begin{aligned} S_{\text{WZ}} &= \mp \pi R \mathcal{T}_5 \int [C_{(5)} + C_{(3)} \wedge (dt \wedge dz + \alpha dz \wedge dv \\ &\quad + h dv \wedge dw) + h C_{(1)} \wedge dt \wedge dz \wedge dv \wedge dw], \end{aligned} \quad (33)$$

where the terms containing a R-R form wedged to dt are irrelevant. The resulting configuration consists of the following objects. First we have a D4-brane stretched along $R^2 \times T^2$ with coordinates (z, ρ, v, w) . Its RR-charge reads

$$\mathcal{T}_4 = \pi R \mathcal{T}_5, \quad (34)$$

which is precisely the relation one would expect when the codimension-one solution in the world volume theory of an unstable Dp -brane on $R^2 \times S^3$ is identified as a BPS $D(p-1)$ -brane on $R^2 \times T^2$. We also have two D2-branes with charges per unit area, $\pi R \mathcal{T}_5 \alpha$ and $\pi R \mathcal{T}_5 h$, which are spanned by the world volume coordinates (t, ρ, w) and (t, ρ, z) , respectively. In addition, there are fundamental strings with flux (28) on cylinder $R \times S^1$ of (z, w) .

In order to study the BPS nature of the solution, we now investigate the energy-momentum tensor (25). For the solution to be a BPS object, it is required that the stress components vanish in the transverse directions, i.e., $T^{\mu\nu} = 0$ with $\mu, \nu = u, v$. From (25), this is satisfied if $\beta \rightarrow 0$ for which the pressure in the u direction T^{uu} vanishes. [The other limit $\alpha \rightarrow 0$ with finite β is not interesting since $\Sigma(u)$ and local densities blow up.] Also, the off-diagonal stress component between D2-branes T^{zw} should vanish. In the thin limit $\beta \rightarrow 0$ where u is fixed to $\pi R/2$, this dictates that $h = 0$. Then we have only one D2-brane along the directions (ρ, w) and an electric flux along the z direction since Π_w becomes zero. This solution is expected to form a BPS configuration. Indeed, the energy per unit area of coordinates (ρ, w) takes the form

$$\begin{aligned} \frac{\mathcal{E}}{\int d\rho dw} &= \int dzdudv(1 + \alpha^2)\Sigma(u) \\ &= \int dzdudv\Pi_z + \pi R\mathcal{T}_5\alpha \int dzdv \\ &= Q_{F1} + Q_{D2}, \end{aligned} \tag{35}$$

where Q_{F1} is the total charge of fundamental strings along the z direction and Q_{D2} the total charge of the D2-brane stretched along the (ρ, w) direction on the area $\int dzdv$. This is the familiar BPS sum rule which we have met in our previous works when considering solutions such as tachyon kinks and tubes [11,16,17]. Other nonvanishing components of the energy-momentum tensor are T^{tv} whose presence means that the configuration carries angular momentum as in tachyon tubes, T^{zz} corresponding to the fundamental string charge, and T^{ww} the D2-brane charge.

In summary, when $h = 0$, the solution produces a BPS D4D2F1-composite in the thin limit, $\beta \rightarrow 0$. It consists of the D4-brane wrapped on $R^2 \times T^2$, the tubular D2-brane with the coordinates (ρ, w) , and the fundamental strings stretched along the z direction.

B. $T = T(w)$

When we put $\partial_u T = 0$ into the equations of motion (16)–(18), we obtain exactly the same equations as in the previous case if $\partial_u T$ is replaced by $\partial_w T$. The solution can be expressed as

$$\sinh\left(\frac{T(w)}{R}\right) = \pm \left[\sqrt{\left(\frac{\mathcal{T}_5}{\alpha\beta}\right)^2 - 1} \cos\left(\frac{w}{R}\right) \right], \tag{36}$$

where β is again given by (22) with $X = -\alpha^2[1 + (\partial_w T)^2]$.

The energy-momentum tensor (24), however, has a different form since the background fields, (5) and (6), are not symmetric under the exchange of u and w . Its nonvanishing components are

$$\begin{aligned} T^{tt} &= \frac{1 + \alpha^2}{\sin(u/R)} \Sigma(w) + \frac{\beta}{\sin(u/R)} \left[h^2 - 2h \cos\left(\frac{u}{R}\right) \right], \\ T^{zz} &= -\frac{1}{\sin(u/R)} \Sigma(w) - \frac{\beta}{\sin(u/R)} \left[h^2 - 2h \cos\left(\frac{u}{R}\right) \right], \\ T^{uu} &= -\frac{\alpha^2}{\sin(u/R)} \Sigma(w), \quad T^{ww} = -\frac{\beta\alpha^2}{\sin(u/R)}, \\ T^{tv} &= -\frac{\alpha}{\sin(u/R)} \Sigma(w), \quad T^{tw} = \alpha\beta \cot\left(\frac{u}{R}\right), \\ T^{zw} &= \frac{\alpha\beta}{\sin(u/R)} \left[h - \cos\left(\frac{u}{R}\right) \right]. \end{aligned} \tag{37}$$

We also have the electric fluxes

$$\begin{aligned} \Pi_z &= \Sigma(w) + \beta \left[h^2 - 2h \cos\left(\frac{u}{R}\right) \right], \\ \Pi_w &= -\alpha\beta \left[h - \cos\left(\frac{u}{R}\right) \right]. \end{aligned} \tag{38}$$

Comparing with the previous case, we see that there are terms which are not proportional to $\Sigma(w)$. But note that they are all multiplied by β and, hence, all the quantities are proportional to $\Sigma(w)$ in the $\beta \rightarrow 0$ limit as before. Moreover, in this thin limit, h completely disappears from the energy-momentum tensor and the electric flux. This is also true for the Wess-Zumino term (31) since $dT = \partial_w T dw$ kills the terms containing h in (32), which means that the constant magnetic field h does not induce D2-brane charge. From the expression for the energy density in (37), it can be seen that a BPS sum rule similar to (35) holds and is independent of h . Therefore, in the present case, one may say that there is an infinite degeneracy with respect to h which is, different from the previous case, not constrained to be zero in the BPS limit. The BPS object is again a D4D2F1-composite with the world volume direction w replaced by u .

Since the range of the variable w is from 0 to $4\pi R$, $\Sigma(w)$ now reduces to a sum of four delta functions rather than one in the thin limit, i.e.,

$$\Sigma(w) \rightarrow \frac{\pi R\mathcal{T}_5}{\alpha} \sum_{n=0}^3 \delta(w - \pi R/2 - n\pi R), \quad \beta \rightarrow 0. \tag{39}$$

Because of this difference and the different pattern of degeneracy in h , it might seem that the solution in this case describes a different object from that of the previous case. We will, however, show that the difference is actually an artifact of the coordinates. It turns out that the two configurations are identical, with the ways T^2 being embedded in S^3 different.

First, we observe that, if the coordinate w is exchanged with u , the energy-momentum tensor (37) and the electric flux (38) are identical to (25) and (28) of the previous case in the thin limit with $h = 0$. Now we consider the geometry of the solution. With $w = (n + \frac{1}{2})\pi R$, $n = 0, \dots, 3$, the

metric of S^3 reduces to

$$ds_{S^3} = du^2 + \sin^2\left(\frac{u}{R}\right)dv^2 + \left[dw + \cos\left(\frac{u}{R}\right)dv \right]^2$$

$$\rightarrow du^2 + dv^2, \quad (40)$$

which is locally the metric of T^2 . The global topology is determined by examining the range of coordinates. In the Appendix, we verify this by finding an explicit orthogonal coordinate transformation connecting the two solutions. Since the background fields are obviously invariant under the transformation, the solution indeed describes the same object as in the previous section embedded in a different direction in S^3 (in the limit $\beta \rightarrow 0$).

IV. CONCLUSION

In this paper we studied the DBI type effective action of a non-BPS D5-brane in the asymptotic limit of the ten-dimensional lift of the Salam-Sezgin vacuum. In the limit the background approaches $R^{1,4} \times T^2 \times S^3$ and exact tachyon kink solutions were found for the D5-brane on $R^{1,2} \times S^3$. There are two classes of solutions both of which describe BPS D4D2F1-composites on $R^2 \times T^2$, D2 being tubular and wrapped on an S^1 and F1 stretched along one of the flat directions. In one class of the solutions, there is an infinite degeneracy with respect to a constant magnetic field along the direction of the NS-NS field on the three-sphere.

Although the ten-dimensional embedding of the Salam-Sezgin vacuum is a solution of type IIA supergravity, it needs to be seen whether it is also an exact string background. One of the methods to check this is to see whether the background survives higher-order stringy correction terms in the low-energy type-IIA string effective action. However, the background at the asymptotic limit may provide an exact string background, considering that the local geometry is that of the near-horizon limit of the NS5-brane which is known to be exact at the string tree level [9].

ACKNOWLEDGMENTS

This work was supported by Korea Research Foundation Grants (KRF-2002-070-C00025 for C. K.) and is the result of research activities [Astrophysical Research Center for the Structure and Evolution of the Cosmos (ARCSEC)] supported by the Korea Science & Engineering Foundation (Y. K., H. K., and O. K.). C. K. and Y. K. thank the KIAS for their hospitality where part of the work was done.

APPENDIX: EQUIVALENCE BETWEEN THE TWO TYPES OF SOLUTIONS IN BPS LIMIT

Here we demonstrate that under a suitable coordinate transformation the solution $T = T(u)$ of (23) in Sec. III A becomes the solution $T = T(w)$ of (36) in Sec. III B in the thin limit.

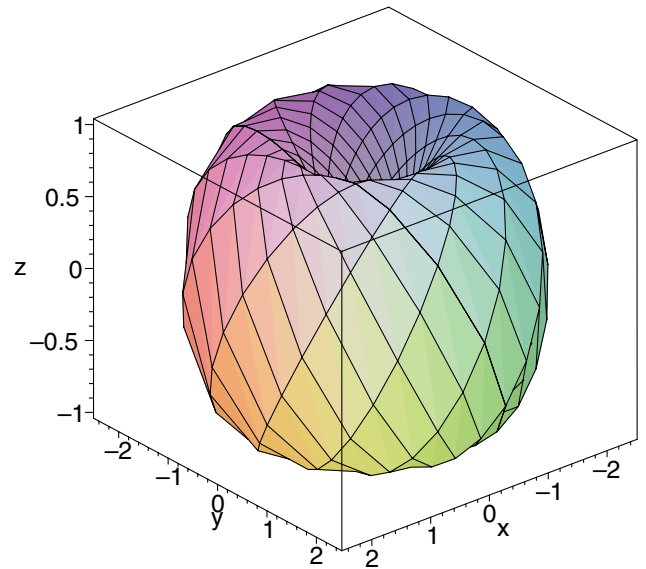


FIG. 1 (color online). Stereographic projection of the solution $T = T(u)$ of Sec. III A.

The three-sphere with metric ds_{S^3} in (30) can be represented by Cartesian coordinates ζ_i , $i = 1, \dots, 4$ via

$$\begin{aligned} \zeta_1 &= \cos\left(\frac{u}{2R}\right) \cos\left(\frac{v+w}{2R}\right), \\ \zeta_2 &= \cos\left(\frac{u}{2R}\right) \sin\left(-\frac{v+w}{2R}\right), \\ \zeta_3 &= \sin\left(\frac{u}{2R}\right) \cos\left(\frac{v-w}{2R}\right), \\ \zeta_4 &= \sin\left(\frac{u}{2R}\right) \sin\left(\frac{v+w}{2R}\right), \end{aligned} \quad (A1)$$

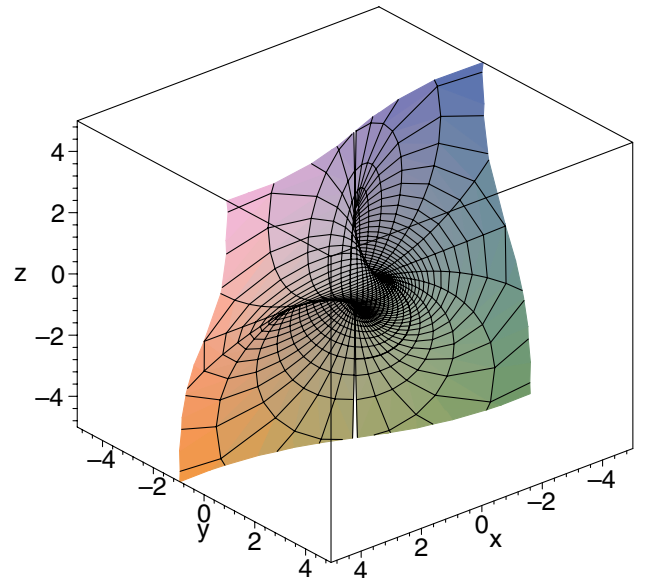


FIG. 2 (color online). Stereographic projection of the solution $T = T(w)$ of Sec. III B.

so that $\sum_{i=1}^4 \zeta_i^2 = 1$. In these coordinates, the solution (23) describes T^2 embedded in S^3 as seen in Sec. III A. To visualize this solution, it is convenient to use the stereographic projection of S^3 onto R^3 given by

$$w_i = \frac{\zeta_i}{1 - \zeta_4} \quad (i = 1, 2, 3), \quad (\text{A2})$$

where the point $(0, 0, 0, 1)$ is chosen to be the ‘‘north pole.’’ Figure 1 shows the torus (30) embedded in S^3 .

In this stereographic projection, the configuration obtained as the thin limit of the solution (36) is visualized as in Fig. 2. Note that the figure is obtained after patching all contributions from the four delta functions in (39). At first,

the shape of the surface in Fig. 2 does not look like a torus. However, the spatial infinities of R^3 are to be identified in the stereographic projection and it is not difficult to see that the surface actually is a torus embedded in S^3 . Indeed, with the new coordinates

$$\begin{aligned} \zeta'_1 &= -\frac{1}{\sqrt{2}}(\zeta_2 + \zeta_4), & \zeta'_2 &= \frac{1}{\sqrt{2}}(\zeta_1 + \zeta_3), \\ \zeta'_3 &= \frac{1}{\sqrt{2}}(\zeta_2 - \zeta_4), & \zeta'_4 &= \frac{1}{\sqrt{2}}(\zeta_1 + \zeta_3), \end{aligned} \quad (\text{A3})$$

and the corresponding stereographic projection, the surface in Fig. 2 is precisely transformed to the torus of Fig. 1.

-
- [1] For a review, see A. Sen, hep-th/0410103, and references therein.
- [2] A. Sen, J. High Energy Phys. 04 (2002) 048; 07 (2002) 065.
- [3] E. Witten, J. High Energy Phys. 12 (1998) 019.
- [4] A. Sen, Mod. Phys. Lett. A **17**, 1797 (2002).
- [5] A. Sen, Phys. Rev. D **68**, 066008 (2003).
- [6] M. Cvetič, G. W. Gibbons, and C. N. Pope, Nucl. Phys. **B677**, 164 (2004).
- [7] A. Salam and E. Sezgin, Phys. Lett. **147B**, 47 (1984).
- [8] D. Kutasov, hep-th/0405058.
- [9] C. G. Callan, J. A. Harvey, and A. Strominger, hep-th/9112030.
- [10] S. J. Rey, Phys. Rev. D **43**, 526 (1991).
- [11] C. Kim, Y. Kim, and C. O. Lee, J. High Energy Phys. 05 (2003) 020; C. Kim, Y. Kim, O. K. Kwon, and C. O. Lee, J. High Energy Phys. 11 (2003) 034.
- [12] T. Okuda and S. Sugimoto, Nucl. Phys. **B647**, 101 (2002).
- [13] N. Lambert, H. Liu, and J. Maldacena, hep-th/0303139.
- [14] M. R. Garousi, Nucl. Phys. **B584**, 284 (2000); E. A. Bergshoeff, M. de Roo, T. C. de Wit, E. Eyras, and S. Panda, J. High Energy Phys. 05 (2000) 009; J. Kluson, Phys. Rev. D **62**, 126003 (2000).
- [15] A. Buchel, P. Langfelder, and J. Walcher, Ann. Phys. (N.Y.) **302**, 78 (2002); C. Kim, H. B. Kim, Y. Kim, and O.-K. Kwon, J. High Energy Phys. 03 (2003) 008.
- [16] C. Kim, Y. Kim, O.-K. Kwon, and P. Yi, J. High Energy Phys. 09 (2003) 042.
- [17] C. Kim, Y. Kim, and O. K. Kwon, J. High Energy Phys. 05 (2004) 020.
- [18] K. Okuyama, J. High Energy Phys. 05 (2003) 005.